

3. RELATIVISTIC WAVE EQUATIONS

- The Schrödinger equation $i\hbar \dot{\Psi} = H\Psi = (\nabla^2 + V)\Psi$ is not relativistically invariant.

x & t have to appear as x^μ , $p^\mu \propto \partial^\mu$ etc.

- Moreover, as $E > m_x + m_{\bar{x}}$, particles - anti-particle pairs are created & number of particles is not conserved, one has to deal with multi-particle states.

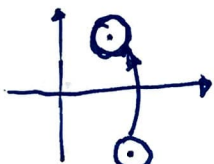
e.g. 

3.1. THE KLEIN-GORDON EQUATION

$\Psi(x^\mu)$... solution, then $\Lambda\Psi$... also a solution

$$\Psi(x^\mu) \rightarrow \Psi'(x^\mu) = \Psi(\Lambda^{-1\mu\nu} x_\nu)$$

\uparrow
canonical (usual) choice



a^μ ... max of $\Psi(x^\mu)$ then $\Lambda a^\mu = a^{\mu'}$

$$\boxed{\Psi'(x') = \Psi(x)} \quad \text{or} \quad \Psi'(x) = \Psi(\Lambda^{-1}x)$$

$$= \Psi(\Lambda^{-1} \underbrace{\Lambda x}_{x'}) = \Psi(x)$$

- If Ψ solves the eqn, then Ψ' should, too.

- Simplest scalar field equation that does this is

$$\partial^\mu \partial_\mu \psi + m^2 \psi = 0 \quad \text{KLEIN-GORDON EQ.}$$

$$\left(\frac{\partial^2}{\partial t^2} - \nabla^2 \right) \psi + m^2 \psi = 0 \Rightarrow \boxed{(\partial^2 + m^2) \psi = 0}$$

$$\partial^\mu \text{ transforms as } x^\mu \quad \partial_\mu = \frac{\partial}{\partial x^\mu}$$

& $\partial^\mu \partial_\mu$ is Lorentz invariant

- What are the solutions to $y'' = -ay$, $y = e^{iax}$

$$\psi(x^\mu) = e^{\pm i p^\mu x_\mu} \quad p^\mu = (E, \vec{p})$$

$$= e^{\pm i (Et - \vec{p} \cdot \vec{x})} \quad \& \quad p^\mu p_\mu = m^2 = \sqrt{E^2 - \vec{p}^2}$$

$$\text{Now: } \psi'(x'^\mu) = \psi(\Lambda^{-1} x'^\mu) = e^{i p^\mu \Lambda^{-1} x'_\mu} = e^{i (\Lambda p)^\mu x'_\mu}$$

because: $x \cdot p = \Lambda x \cdot \Lambda p$ is L.I.

$$\Lambda^{-1} x p = x \cdot \Lambda p = (\Lambda p) \cdot (\Lambda p) = p^2$$

- Note that the on-shell condition has two

$$\text{solutions: } E^2 - \vec{p}^2 = m^2, \quad E = \pm \sqrt{\vec{p}^2 + m^2}$$

- The QFT interprets this as particles & anti-particles.

- A more transparent way to derive the e.o.m. is by starting from a L.I. scalar ^{function = action} and require it to be extremized, similar to Lagrangian in classical mechanics: $S[x, \dot{x}, t] = \int dt L(x, \dot{x})$

- Instead of t only, we need to integrate over the entire 4-space

$$S(\varphi(x^\mu)) = \int d^4x \mathcal{L}(\varphi, \partial_\mu \varphi)$$

↑
↑
 Lorentz-invariant (LI) integral L.I. Action, scalar

consider a single real scalar field then

- $\mathcal{L} = \frac{1}{2} \partial_\mu \varphi \partial^\mu \varphi - \frac{1}{2} m^2 \varphi^2$ is a good candidate. One can add higher order terms, like $\lambda \varphi^4$ to get interactions.

- Extremizing S gives $\delta S = \frac{\delta \mathcal{L}}{\delta \varphi} \delta \varphi + \frac{\delta \mathcal{L}}{\delta \partial_\mu \varphi} \delta \partial_\mu \varphi$

$$= \delta \varphi \left(\frac{\delta \mathcal{L}}{\delta \varphi} - \partial_\mu \frac{\delta \mathcal{L}}{\delta \partial_\mu \varphi} \right) + \partial_\mu \left(\frac{\delta \mathcal{L}}{\delta \partial_\mu \varphi} \delta \varphi \right) \underset{=0}{=} 0$$

E.o.m: $\partial_\mu \frac{\delta \mathcal{L}}{\delta \partial_\mu \varphi} - \frac{\delta \mathcal{L}}{\delta \varphi} = 0$

$\Rightarrow \partial_\mu \partial^\mu \varphi + m^2 \varphi = 0$ (ok)

- Because we extremized the action $S[\psi]$ by finding the solution, and because S is LI, it follows that $S[\psi'] = S[\psi]$ and ψ' also extremizes S and thereby solves the equations of motion.

3.2 FIELDS AND PARTICLES

- We have to associate the field operator $\psi(x^\mu)$ to the particles that we observe (their charges, spin, momentum, energy, mass and other properties).
- In QFT this is done by introducing a set of states (multiparticle) called the Fock space.

e.g. $|0\rangle$... vacuum, empty state, no particles

$|\psi(p)\rangle$... state that represents a single particle with momentum p

$\langle \text{BRA} | \text{OPERATOR} | \text{KET} \rangle$... function, solution

$$\langle 0 | \psi(x^\mu) | \psi(p) \rangle = e^{-i p x}$$

final state,
creates a vacuum

field operator at x^μ

-93-4-

initial state
to be annihilated

BRA = FINAL STATE FIELD OPERATOR KET = INITIAL STATE

$$\underbrace{\langle 0} \underbrace{| \varphi(x^\mu)} \underbrace{| \varphi(p) \rangle} = e^{-ipx}$$

vacuum state, empty - no particles real scalar particle state with momentum p wavefunction of the annihilated scalar particle w. momentum p & $E_p^2 - p^2 = m^2$

real scalar field operator acting at x^μ

this operator annihilates $|\varphi(p)\rangle$ into the vacuum

* taking the charge conjugated equation (this is a real scalar field) gives

$$\langle \varphi(p) | \varphi^\dagger(x^\mu) | 0 \rangle = e^{+ipx}$$

* Here, the final state is the bosonic particle with momentum p . It was created from the vacuum by the creation operator $\varphi^\dagger(x^\mu)$ at the position x^μ from the initial vacuum state $|0\rangle$. The RHS is φ^* of the particle when it was created, it has $E = E_p = \sqrt{p^2 + m^2}$ & mass m .

* If we are dealing with a complex scalar field

$$\mathcal{L} = \partial_\mu \varphi (\partial^\mu \varphi)^\dagger - m^2 \varphi^\dagger \varphi, \quad \varphi = \frac{1}{\sqrt{2}}(\varphi_1 + i\varphi_2)$$

Then we have $\psi^+ \neq \psi$ ↙ anti-particle state

$$\langle 0 | \psi^+ | \psi(p) \rangle = e^{-ipx}$$

$$\langle 0 | \psi | \psi(p) \rangle = e^{-ipx}$$

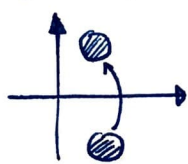
3.3. MAXWELL'S EQUATIONS

- Scalar fields have masses & carry charges. How do we describe gauge bosons with spin 1 & fermions with spin $\frac{1}{2}$?

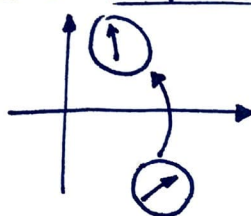
SPIN 1 : Photons, Weak bosons W, Z , mesons, like ρ^0, \dots

- This is the vector representation of the rotation group

SCALAR



VECTOR & FERMION



- Vectors are $v^i, v_i, \Rightarrow x^\mu, p^\mu$ & transform

as
$$v^i = R^{ij} v_j \quad \text{or} \quad x^{\mu'} = \Lambda^{\mu\nu} x_\nu$$

so the field goes as
$$A^{\mu'}(x^{\mu'}) = \Lambda A^\mu(x)$$

in short:
$$A' = \Lambda A(\Lambda^{-1}x)$$

- So we rotate the field & transform the argument

$$A'(x) = \Lambda A(\Lambda^{-1}x)$$

(in general we have one Λ for each index, i.e. 0 for the scalar field)

- As for the scalar field, we have the states that characterize the mass, momentum & now also spin \rightarrow polarization ϵ . 3 polarizations
↓

• 3D: $\langle \underbrace{0}_{\text{vacuum}} | \underbrace{V^i(x)}_{\text{gauge boson, or vector operator for each polarization}} | \underbrace{\psi(p, \epsilon)}_{\text{vector particle that gets annihilated}} \rangle = \epsilon^i e^{-ipx}$
↑
solves the K.G. eqn

- of course, the ϵ^i is a vector that also rotates
 $\epsilon'^i = R^{ij} \epsilon^j$
- higher spins have more degrees of freedom related

• to spin s , i.e. $2s+1$ -- degeneracy factor = # of degrees of freedom so R generalizes $R^{(s)ij}$

- generalization to Minkowski space-time in 4D has to be done carefully because $\mu=0,1,2,3 \dots$ 4D & $s=1$, thus $2s+1=3 < 4$, which means that we have too many d.o.f. s in ϵ^μ than needed.

- Let us generalize to 4D by the ansatz

$$\langle 0 | V^\mu(x) | \nu(p, \epsilon) \rangle = \epsilon^\mu e^{-ipx}$$

- Now ϵ^μ is a 4-vector that rotates under Lorentz transformations, as usual $\epsilon'^\mu = \Lambda^\mu_\nu \epsilon^\nu$

- For momenta we have $(E, \vec{p}) = p^\mu$ & we

- normalize to $p^\mu p_\mu = E^2 - \vec{p}^2 = m^2 > 0$

- For polarization vectors we have

$$\epsilon^{(0)\mu} = (1, 0, 0, 0) \quad \text{time-like pol. mode}$$

$$\epsilon^{(\lambda)\mu} = (0, \vec{\epsilon}^{(\lambda)}(p)) \quad \lambda=1,2 \dots \text{transverse modes}$$

- we can align $\vec{p} = (E, 0, 0, p)$ with $\lambda=3 \dots$ longitudinal mode

$$\Rightarrow p^\mu \epsilon_{\mu}^{(\lambda)} = 0 \quad \Rightarrow \text{transverse modes}$$

- Normalization for ϵ_μ

$$(\lambda=0) \quad \epsilon_\mu^{(0)} \epsilon^{\mu(0)} = \epsilon^{(0)} \cdot \epsilon^{(0)} = -1$$

$$\lambda=1,2,3 \quad \epsilon_\mu^{(\lambda)} \epsilon^{\mu(\lambda)} = 1$$

- Sum over polarizations

$$\sum_{\lambda} \epsilon_\mu^{(\lambda)} \epsilon_\nu^{(\lambda)} = -g_{\mu\nu}$$

for a $m=0$
gauge boson,
massless photon

- The timelike $\epsilon_{\mu}^{(0)} = (1, 0, 0, 0)$ is unphysical in a sense that it does not affect the EM interactions

- 3D Maxwell's equations give $\vec{\epsilon}$ in the solution

$$\vec{A}(x) = \underbrace{\vec{\epsilon} e^{-iEt + ip\vec{x}}}_{\text{this is the wave-function of } \vec{A}, \text{ the photon}}$$

$$p^\mu = (E, 0, 0, E) \Rightarrow E = p$$

- we also know that $\vec{E} \perp \vec{B} + \vec{p} \Rightarrow \underline{\underline{\vec{\epsilon} \cdot \vec{p} = 0}}$

- As we saw, the p_x remains unchanged under

$$L_{\pi} \Rightarrow \Lambda_p \Lambda p = x p$$

$$\times \Lambda p = \Lambda^{-1} x p$$

- but e.g. $\epsilon = (0, \epsilon^1, \epsilon^2, 0) \rightarrow (\#, \#, \epsilon^2, 0)$

under a boost in the x -direction. This timelike component has no impact on the interactions, as long as the em current $\bar{\psi} \gamma^\mu \psi = j^{\mu em}$ is conserved: $\partial_\mu j^\mu = 0$

- For the massive g.b. fields one has additional fields (Goldstone bosons) that enter in the longitudinal modes.

- As for the scalars, we can cast the Maxwell's eqn in a variational form by defining

$$A^\mu = (\phi, \vec{A}), \quad F^{\mu\nu} = \partial^\mu A^\nu - \partial^\nu A^\mu$$

↓
4D E.M. vector potential

↓
relativistic anti-symmetric E.M. tensor, field strength tensor

$$x^\mu = (ct, \vec{x}) \quad \partial^\mu = \frac{\partial}{\partial x^\mu} = \left(\frac{d}{cdt}, -\nabla \right)^\mu$$

$$\Rightarrow F^{i0} = (-\nabla^i \phi - \partial_0 A^i) = E^i \text{ electric field}$$

$$\mathbb{F}^{ij} = (-\nabla^i A^j + \nabla^j A^i) = -\epsilon^{ijk} B^k \text{ magnetic field}$$

- By contracting the Lorentz indices, we have

$$\mathcal{L} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} - j^\mu A_\mu$$

↙
canonical normalization

↖
Lorentz scalar

↗
any external current, e.g.

$\psi \gamma^\mu \psi$, that satisfies

$$j^\mu = (\rho, \vec{j})$$

↙
charge distribution

↖
current density

$$\partial_0 \rho - \nabla \cdot \vec{j} = 0$$

$$\underline{\partial_\mu j^\mu = 0.}$$

- By requiring $\delta \mathcal{L} = 0$ we extremize the action and get the coupled e.o.m. (i.e. Maxwell's eqs.)

$$\begin{aligned}
\mathcal{L} &= -\frac{1}{4} (\partial^\mu A^\nu - \partial^\nu A^\mu) (\partial_\mu A_\nu - \partial_\nu A_\mu) - A_\mu j^\mu / \delta_{\text{on}} A \\
&= -\frac{1}{2} (\partial_\mu \delta A_\nu - \partial_\nu \delta A_\mu) F^{\mu\nu} - \delta A_\mu j^\mu = 0 \\
&= -\frac{1}{2} \delta A_\nu \left\{ \partial_\mu F^{\mu\nu} + \frac{1}{2} \partial_\nu \delta A_\mu F^{\mu\nu} - \delta A_\nu j^\mu \right\} \\
&= +\frac{1}{2} \delta A_\nu (\partial_\mu F^{\mu\nu} + \partial_\mu F^{\mu\nu}) - \delta A_\nu j^\nu \\
&= \delta A_\nu (\partial_\mu F^{\mu\nu} - j^\nu) = 0
\end{aligned}$$

• \Rightarrow $\boxed{\partial_\mu F^{\mu\nu} = j^\nu}$; if we insert the $\begin{pmatrix} 0 & E_x & E_y & \dots \end{pmatrix}$

we get: $\vec{\nabla} \cdot \vec{E} = j$, $\epsilon^{ijk} \partial_j B^k = \frac{d}{dt} E^i = j^i$

3.4 THE DIRAC EQUATION

- spin $\frac{1}{2}$ electron needs $2s+1 = 2$ states, right?
- Which Lorentz representation behaves like that and how do we write down the e.o.u. / action and how does it couple to gauge bosons A^μ & scalars ϕ ?
- Dirac introduced a set of 4 matrices γ^μ that, when combined with the spinorial representation ψ , transforms as a Lorentz vector,

$$\bar{\psi} \gamma^\mu \psi \rightarrow \Lambda^\mu_\nu \bar{\psi} \gamma^\nu \psi.$$

- These matrices obey the Clifford algebra, which is the anti-commuting Minkowskian implementation

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu} = 2 \begin{pmatrix} 1 & & & \\ & -1 & & \\ & & -1 & \\ & & & -1 \end{pmatrix}$$

- There are infinitely many representations (actual numbers) of these matrices, connected by LT. These two are ~~not~~ very commonly used:

$$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \quad \gamma^i = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} \quad \text{NON-RELATIVISTIC}$$

$$\gamma^\mu = \begin{pmatrix} 0 & \sigma^\mu \\ \bar{\sigma}^\mu & 0 \end{pmatrix} \quad \sigma^\mu = (1, \sigma^i) \quad \text{RELATIVISTIC} \\ \bar{\sigma}^\mu = (1, -\sigma^i) \quad \text{CHIRAL REP.}$$

- of course $\sigma^1 = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}$, $\sigma^2 = \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}$, $\sigma^3 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$ are the usual Pauli matrices.

- We can try to contract γ^μ with ∂_μ to build a LI object and solve its e.o.m.

DIRAC EQUATION $\rightarrow (i \gamma^\mu \partial_\mu - m) \psi = 0$

↑
four-component Dirac spinor

- As the scalars & gauge boson, these objects obey the K.G. equation such that $\psi \sim \xi^s \cdot e^{-ipx}$

• We will see that ψ :
 • i) obeys the K.G. eq.

• ii) gives a conserved current $\bar{\psi}\gamma^0\psi$

• iii) has $E_p > 0$ & $E_p < 0$ solutions when

$\vec{p}=0$ (at rest) : particles & anti-particles

i) To see that ψ obeys the KG eq. let's take the

"absolute value" of the Dirac equation

$$(i\gamma^\mu \partial_\mu - m)\psi = 0 \quad / \quad (i\gamma^\mu \partial_\mu + m)$$

$$= (i^2 \gamma^\mu \partial_\mu \gamma^\nu \partial_\nu \pm m^2)\psi$$

$$= \left(-\frac{1}{2} \underbrace{(\gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu)}_{2g^{\mu\nu}} \partial_\mu \partial_\nu - m^2 \right) \psi = - \underbrace{(\partial_\mu \partial^\mu + m^2)}_{\text{KLEIN-GORDON EQ. QED.}} \psi = 0$$

ii) To analyze the solution at rest, let's go to the rest frame with $\vec{p}=0$, where $e^{-ipx} \rightarrow e^{-iEt}$

and, as we said, the NON-REL. representation is

useful $\psi = \begin{pmatrix} a \\ b \\ c \\ d \end{pmatrix} e^{-iEt}$, $\gamma^0 E - \gamma^i \vec{p}_i = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} E$

$$\Rightarrow (\pm \gamma^0 E - m)\psi = \begin{pmatrix} E-m & & & 0 \\ & E-m & & \\ & & -E-m & \\ 0 & & & -E-m \end{pmatrix} \begin{pmatrix} a \\ b \\ c \\ d \end{pmatrix} e^{-iEt} = 0$$

a) $E=m$ $\psi_+ = \begin{pmatrix} a \\ b \\ 0 \\ 0 \end{pmatrix} e^{-iEt}$, $\psi_- = \begin{pmatrix} 0 \\ 0 \\ c \\ d \end{pmatrix} e^{-iEt}$

- Clearly, we have two disjoint solutions

$$\psi_+ = \begin{pmatrix} \xi \\ 0 \end{pmatrix} e^{-imt}, \quad \psi_- = \begin{pmatrix} 0 \\ \eta \end{pmatrix} e^{+imt}$$

for a massive Dirac fermion at rest.
PARTICLES (Electrons) ANTI-PARTICLES (positrons)

- We can perform a Lorentz boost from the rest frame and find a general solution

$$(i\gamma^\mu \partial_\mu - m)\psi = \begin{pmatrix} E-m & -\vec{\sigma} \cdot \vec{p} \\ \vec{\sigma} \cdot \vec{p} & -E+m \end{pmatrix} \begin{pmatrix} \alpha \\ \beta \\ 0 \\ d \end{pmatrix} e^{-iEt + \vec{p} \cdot \vec{x}} = 0.$$

- The positive energy solution is $\psi_+ = u^{(s)}(p) e^{-ipx}$

where $s=1,2$ is the spin index

$$\xi = \begin{pmatrix} 1 \\ 0 \end{pmatrix} \dots \text{spin up } e^- \quad \eta = \begin{pmatrix} 0 \\ 1 \end{pmatrix} \dots \text{spin down, etc.}$$

$u^{(s)}(p) \dots$ particle spin orientation

$v^{(s)}(p) \dots$ anti-particle

EXAMPLE

$$\psi_+ = \begin{pmatrix} \xi \\ 0 \end{pmatrix} e^{-imt} \xrightarrow{\text{BOOST}} \begin{pmatrix} \xi \\ -\frac{\vec{\sigma} \cdot \vec{p}}{2m} \xi \end{pmatrix} e^{-iE_p t + i\vec{p} \cdot \vec{x}}$$

$\sqrt{p^2 + m^2}$
 $+ O(p^2)$

- Now we can introduce the proper Fock states.

• FERMIONIC STATES & OPERATORS

the wave function of the annihilated $e^-(s,p)$

$$\langle \underbrace{0}_{\text{vacuum}} | \underbrace{\psi(x)}_{\substack{\uparrow \\ \text{annihilates an} \\ e^- \text{ at } x^\mu}} | \underbrace{e^-(p,s)}_{\substack{\text{electron} \\ \text{state with} \\ \text{momentum} \\ p \text{ \& spins}}} \rangle = \underbrace{u^{(s)}(p)}_{\substack{\text{spinorial} \\ \text{part} \sim \\ \text{polarization}}} \underbrace{e^{-ipx}}_{\substack{\text{the usual} \\ \text{plane-wave} \\ \text{part}}}$$

$$p^\mu = (E_r, \vec{p})$$

The conjugate part represents the creation of the e^-

$$\langle e^-(p,s) | \psi^\dagger(x) | 0 \rangle = u^{(s)\dagger}(p) e^{+ipx}$$

• Similarly, we have creation & annihilation operators for the anti-particle, by convention:

$$\langle 0 | \psi^\dagger(x) | e^+(p,s) \rangle = \underline{v}^{(s)\dagger}(p) e^{-ipx}$$

$$\langle e^+(p,s) | \psi(x) | 0 \rangle = v^{(s)}(p) e^{ipx}$$

iii) CONSERVED GAUGE CURRENT j^μ FROM FERMIONS

$$\bullet (\gamma^\mu)^\dagger = ? \quad (\gamma^0)^\dagger = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}^T = \gamma^0 \quad \text{anti-Hermitian}$$

$$(\gamma^i)^\dagger = \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix}^\dagger = \begin{pmatrix} 0 & -\sigma^{i\dagger} \\ \sigma^{i\dagger} & 0 \end{pmatrix} = -\gamma^i$$

• note that $\gamma^0 \gamma^i = -\gamma^i \gamma^0$

$$\begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} = \begin{pmatrix} -\sigma^i & 0 \\ 0 & \sigma^i \end{pmatrix} \quad - \begin{pmatrix} 0 & \sigma^i \\ -\sigma^i & 0 \end{pmatrix} \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} = - \begin{pmatrix} \sigma^i & 0 \\ 0 & -\sigma^i \end{pmatrix}$$

- The $\gamma^0 \gamma^i = -\gamma^i \gamma^0$ identity means that we can form $(\gamma^0 \gamma^\mu)$ which is Hermitian, i.e.

$$(\gamma^0 \gamma^\mu)^\dagger = \gamma^{\mu\dagger} \gamma^{0\dagger} = \begin{cases} \gamma^0 \gamma^0 = \gamma^0 \gamma^0 \\ -\gamma^i \gamma^0 = \gamma^0 \gamma^i \end{cases} = \gamma^0 \gamma^\mu$$

- Therefore $\psi^\dagger (\gamma^0 \gamma^\mu) \psi$ will also be Hermitian

$$(\psi^\dagger (\gamma^0 \gamma^\mu) \psi)^\dagger = \psi^\dagger (\gamma^0 \gamma^\mu)^\dagger \psi = \psi^\dagger \gamma^0 \gamma^\mu \psi$$

- For convenience we call $\psi^\dagger \gamma^0 = \bar{\psi}$ such that

$\bar{\psi} \gamma^\mu \psi$ is a Hermitian object, a good candidate for $j^\mu = \bar{\psi} \gamma^\mu \psi$, for which $\partial_\mu j^\mu = 0$

- Let's work out the e.o.m. for this current

$$i \gamma^\mu \partial_\mu \psi - m \psi = 0 \quad |^\dagger$$

$$-i \partial_\mu \psi^\dagger \gamma^{\mu\dagger} - m \psi^\dagger = 0 \quad | \cdot \gamma^0$$

$$-i \partial_\mu (\psi^\dagger \gamma^0) \gamma^\mu - m \psi^\dagger \gamma^0 = -i (\partial_\mu \bar{\psi}) \gamma^\mu - m \bar{\psi} = 0$$

such that

$$-i \partial_\mu (\bar{\psi} \gamma^\mu \psi) = -i \partial_\mu \bar{\psi} \gamma^\mu \psi - i \bar{\psi} \gamma^\mu \partial_\mu \psi$$

$$= m \bar{\psi} \gamma^\mu \psi - m \bar{\psi} \gamma^\mu \psi = 0. \quad \text{very good.}$$

- Remember how we introduced the coupling to the EM field in non-relativistic QM

$$\hat{p} \rightarrow \hat{p} + e \hat{A}$$

or $\partial_\mu \rightarrow \partial_\mu + ie A_\mu$, therefore the

Dirac equation becomes

$$(i\gamma^\mu \partial_\mu - m)\psi = 0$$

\Downarrow

$$(i\gamma^\mu D_\mu - m)\psi, \quad D_\mu = \partial_\mu + ie A_\mu$$

covariant derivative

The Lagrangian should also be modified

FREE DIRAC $\mathcal{L} = \bar{\psi}(i\not{D} - m)\psi$

QED: FERMION & A_μ : $\mathcal{L} = \underbrace{-\frac{1}{4} F_{\mu\nu} F^{\mu\nu}}_{\text{MAXWELL'S EQUATIONS}} + \underbrace{\bar{\psi}(i\not{D} - m)\psi}_{\text{FREE FERMION \& EM CURRENT}}$

MAXWELL'S EQUATIONS

FREE FERMION & EM CURRENT

$$\mathcal{L} \ni i(i e A_\mu \bar{\psi} \gamma^\mu \psi) = -e A_\mu \underbrace{[e \bar{\psi} \gamma^\mu \psi]}_{j^\mu}$$

So we identified the QED four-current j^μ

from fermions: $j^\mu = e \bar{\psi} \gamma^\mu \psi$

3.5. RELATIVISTIC NORMALIZATION OF STATES

Our equations of motion are LI because we derived them from the action, which is a Lorentz scalar.

To describe physical processes (decays and scatterings) we have to introduce the states in a LI way.

• NR QM : $\langle p_1 | p_2 \rangle = (2\pi)^3 \delta^{(3)}(\vec{p}_1 - \vec{p}_2)$

This is not invariant under LT, e.g. under a boost in the z -direction. However

REL : $\langle p_1 | p_2 \rangle = 2 E_{p_1} (2\pi)^3 \delta^{(3)}(\vec{p}_1 - \vec{p}_2)$

• Let's see in what sense this definition is LI.

$$\begin{pmatrix} E_p' \\ p^{3'} \end{pmatrix} = \begin{pmatrix} \gamma & \gamma\beta \\ \gamma\beta & \gamma \end{pmatrix} \begin{pmatrix} E_p \\ p^3 \end{pmatrix} = \begin{pmatrix} \gamma E_p + \gamma\beta p^3 \\ \gamma p^3 + \gamma\beta E_p \end{pmatrix}$$

Let's check :

$$\begin{aligned} \langle p_1' | p_2' \rangle &= 2 E_{p_1'} (2\pi)^3 \delta^{(3)}(\vec{p}_1' - \vec{p}_2') \\ &= 2 \gamma (E_{p_1} + \beta p_1^3) (2\pi)^3 \delta(\gamma(p_1^3 + \beta E_{p_1}) - p_2^{3'}) \\ &\quad \delta(p_1^{1'} - p_2^{1'}) \delta(p_1^{2'} - p_2^{2'}) \end{aligned}$$

- Relativistic normalization of states

$$\langle p_1 | p_2 \rangle = 2 E_{p_1} (2\pi)^3 \delta^{(3)}(\vec{p}_1 - \vec{p}_2)$$

- Boost along the 3-direction: $E_p' = \gamma (E_p + \beta p^3)$
 $p_1^0 = p_1^0$, $p_2^0 = p_2^0$
 $p_3^0 = \gamma (p_3^0 + \beta E_p)$

- The normalization then becomes

$$\begin{aligned} \langle p_1' | p_2' \rangle &= 2 E_{p_1'} (2\pi)^3 \delta^{(3)}(\vec{p}_1' - \vec{p}_2') \\ &= 2 \gamma (E_p + \beta p^3) (2\pi)^3 \delta(p_1^{1'} - p_2^{1'}) \left. \begin{array}{l} \delta(p_1^{2'} - p_2^{2'}) \\ \delta(p_1^{3'} - p_2^{3'}) \end{array} \right\} \delta(p_1^0 - p_2^0) \\ &\quad \delta(\gamma(p_1^3 + \beta E_{p_1}) - p_2^{3'}) \end{aligned}$$

- The δ function of an arbitrary $g(x)$ is given by

$$\delta(g(x)) = \frac{1}{|g'(x)|} \quad E_{p_1}^{\cancel{2}} = \sqrt{p_1^2 + m^2}$$

$$\delta(g(p_1^3)) = \frac{1}{\gamma + \gamma \beta \frac{dE_{p_1}}{dp_1^3}} \delta(p_1^3 - p_2^3) \quad \frac{dE_{p_1}}{dp_1^3} = \frac{1}{\cancel{E_{p_1}}} \cdot \cancel{p_1^3}$$

$$= \frac{E_{p_1}}{\gamma (E_{p_1} + \beta p_1^3)} \delta(p_1^3 - p_2^3) = \frac{E_{p_1}}{E_{p_1}'} \delta(p_1^3 - p_2^3)$$

$$\Rightarrow \langle p_1' | p_2' \rangle = 2 \cancel{E_{p_1}} (2\pi)^3 \frac{E_{p_1}}{\cancel{E_{p_1}}} \delta^{(3)}(\vec{p}_1 - \vec{p}_2) = 2 E_{p_1} (2\pi)^3 \delta^{(3)}(\vec{p}_1 - \vec{p}_2)$$

- Very good, this is indeed Lorentz invariant. Let's check the overall $-p_3-19-$ normalization.

• PHASE SPACE INTEGRAL

$\int \frac{d^3 p}{(2\pi)^3}$... this is $SO(3)$ invariant but not LI.

$\int \frac{d^4 p}{(2\pi)^4} 2\pi \delta(p^2 - m^2)$
 ↗ on-shell condition

$$\int dp^0 \delta(p^2 - m^2) = \int dp^0 (\delta(p^{0^2} - \vec{p}^2 - m^2))$$

$$= \frac{1}{|2p^0|} \delta(p^0 - E_p) = \frac{1}{2E_p}$$

$$\int \frac{d^4 p}{(2\pi)^4} 2\pi \delta(p^2 - m^2) = \int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p}$$

• The sum over all states becomes a properly relativistically normalized integral

$$\int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} |p\rangle \langle p| = 1$$

• When applied to a particular state, we have

$$\int \frac{d^3 p}{(2\pi)^3} \frac{1}{2E_p} |p\rangle \langle p| k\rangle = \int d^3 p \delta^{(3)}(\vec{p} - \vec{k}) |p\rangle \langle p| k\rangle$$

$$= \frac{1}{2E_p (2\pi)^3} \delta^{(3)}(\vec{p} - \vec{k}) = |k\rangle.$$

• since $\langle p_1 | p_2 \rangle = 2E_{p_1} \delta^{(3)}(\vec{p}_1 - \vec{p}_2) (2\pi)^3$

$$2 [|p\rangle] = 1 - 3 = -2 \Rightarrow [|p\rangle] = -1 \text{ or } \text{GeV}^{-1}$$

3.6. SPIN & STATISTICS THEOREM / CONNECTION

$| \rangle \dots$ states carry information on spin.

In QFT this information is brought in by the appropriate quantization conditions, required by unitarity & positivity. (Dirac, Feynman, Wightman & Streater)

$S \in \mathbb{N} \Rightarrow$ BOSE-EINSTEIN STATISTICS

$S \in \frac{\mathbb{N}}{2} \Rightarrow$ FERMION-DIRAC STATISTICS